

# Detection of NO<sub>2</sub> emissions using MIPAS



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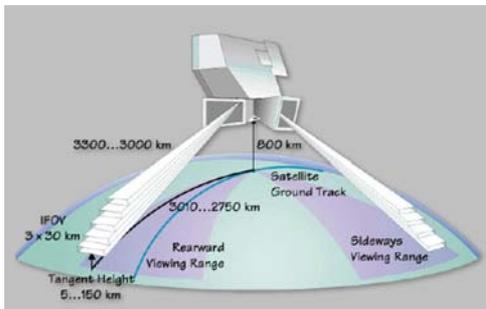
## INTRODUCTION

The Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) is a limb-viewing infrared fourier transform interferometers on the ESA Envisat satellite launched in 2002.

MIPAS makes global measurements of infrared spectra in the range 4-15 $\mu$ m at tangent heights from the mesosphere down to the mid-troposphere. From these spectra, vertical profiles of atmospheric temperature and composition, including stratospheric NO<sub>2</sub>, are routinely retrieved by ESA and several other groups, including Oxford.

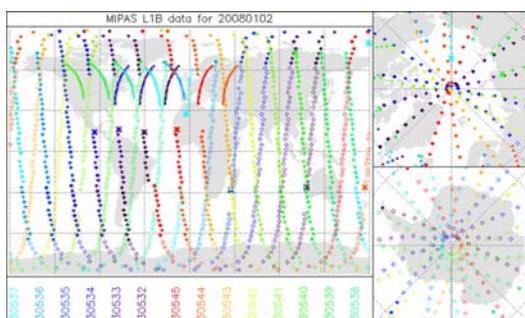
Here we examine the feasibility of retrieving NO<sub>2</sub> concentrations in the upper troposphere from MIPAS spectra.

## THE MIPAS INSTRUMENT

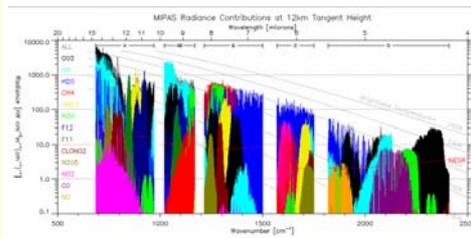


In normal operation MIPAS views rearward along the orbit track giving pole-to-pole coverage and approximately 14 orbits per day. However it also has the ability to view sideways, and such observations are occasionally scheduled to view east-west along the North Atlantic flight corridor as a special mode dedicated to detection of aircraft emissions. Such observations are not yet routinely processed.

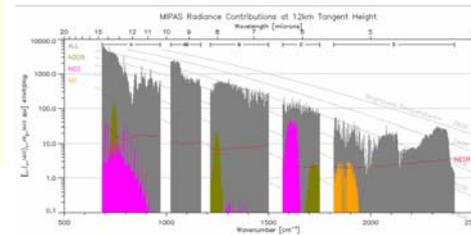
The plot below shows the coverage on 2nd January 2008, a day when such 'AE mode' observations can be seen as breaks from the normal observations following the orbit track.



## THE MIPAS SPECTRUM



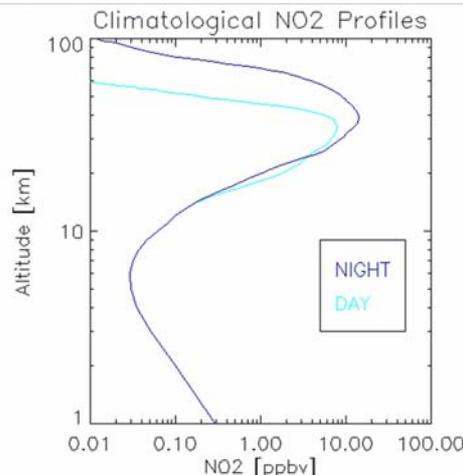
Above is a simulated MIPAS spectrum for a tangent height of 12km showing the spectral features associated with the main infrared emitting species in the atmosphere. Below is the same plot but just showing the NO<sub>x</sub> species.



The grey spectrum represents the total spectrum contributed by all emitting molecules and the red line the MIPAS noise level. On this basis, the NO<sub>2</sub> feature between 1550-1650 cm<sup>-1</sup> has been selected as the most likely to yield a detectable aircraft emission signature.

## ATMOSPHERIC NO<sub>2</sub>

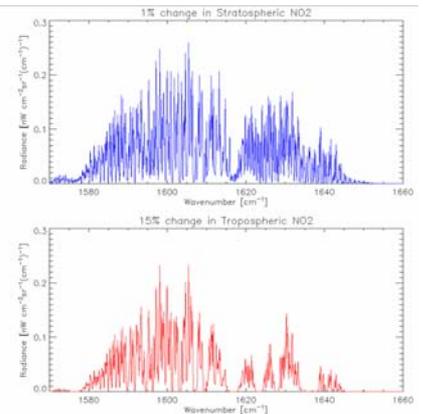
NO<sub>2</sub> (and other NO<sub>x</sub> species) occur naturally in the atmosphere, and have maximum concentrations (measured in ppbv) in the mid-stratosphere.



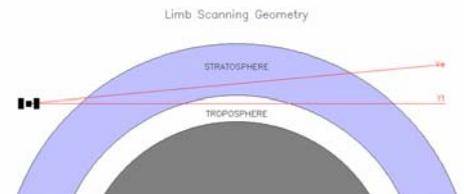
This poses a problem for detecting any enhancement in the upper troposphere from space: it has to be distinguishable from the relatively large stratospheric NO<sub>2</sub> which will also contribute to the measurement.

## NO<sub>2</sub> SENSITIVITY

The plots below show the expected impact of variations in NO<sub>2</sub> in the stratosphere and upper troposphere in the spectrum taken from a 12km tangent height (note the different perturbation sizes)



As expected the two signatures are very similar, making it difficult to assign any enhancement in the signal to a particular altitude along the limb path. However, with limb-scanning instruments such as MIPAS, additional measurements are also available which depend only on the stratospheric column, as shown below.



## CONCLUSIONS

Assuming that the full NO<sub>2</sub> band is used and the above viewing geometry, a linear least-squares-fit analysis suggests that the NO<sub>2</sub> at 12km can be retrieved with a **precision of 3%** (i.e. that an aircraft-induced enhancement of 3% above background levels should be detectable).

However there are some important caveats

- ❖ No other errors, e.g. in temperature or interfering species (primarily H<sub>2</sub>O)
- ❖ The 'shape' of the stratospheric NO<sub>2</sub> is assumed fixed and known
- ❖ In practice, using the entire NO<sub>2</sub> band would be computationally prohibitive. (Operational retrievals tend to concentrate on a few 'microwindows' totalling 10-20cm<sup>-1</sup> in bandwidth).