

# A

## *Tables of Constants*

**TABLE A.1**

Physical Constants

Avogadro's number	$N_a$	$6.022\,141\,79 \times 10^{23}$ molecule mol <sup>-1</sup>
Boltzmann's constant	$k_B$	1.380 650 4 J K <sup>-1</sup>
Electron charge	$e$	$1.602\,176\,487 \times 10^{-19}$ C
Gravitational constant	$G$	$6.674\,29 \times 10^{-11}$ m <sup>3</sup> kg <sup>-1</sup> s <sup>-2</sup>
Magnetic permeability of free space	$\mu_0$	$4\pi \times 10^{-7}$ kg m s <sup>-2</sup> A <sup>-2</sup>
Mass of an electron	$m_e$	$9.109\,382\,15 \times 10^{-31}$ kg
Permittivity of free space	$\epsilon_0$	$8.854\,187\,817 \times 10^{-12}$ kg <sup>-1</sup> m <sup>-3</sup> s <sup>4</sup> A <sup>2</sup>
Planck's constant	$h$	$6.626\,068\,96 \times 10^{-34}$ J s
Stefan-Boltzmann constant	$\sigma$	$5.670\,400 \times 10^{-8}$ J m <sup>-2</sup> s <sup>-1</sup> K <sup>-4</sup>
Speed of light in a vacuum	$c$	$2.997\,924\,58 \times 10^8$ m s <sup>-1</sup>
Universal gas constant	$R$	8.314 472 J mol <sup>-1</sup> K <sup>-1</sup>

**TABLE A.2**  
Astronomical Values

Acceleration of gravity (at sea level and 45 °latitude)	$g$	$9.80616 \text{ m s}^{-2}$
Standard gravity (nominal global average)	$g$	$9.80665 \text{ m s}^{-2}$
Angular velocity of rotation of the Earth	$\omega$	$7.27221 \times 10^5 \text{ rad s}^{-1}$
Average distance, Sun to Earth	$D_S$	$1.496 \times 10^8 \text{ km}$
Average distance, Earth to Moon	$D_M$	$3.84 \times 10^5 \text{ km}$
Radius of the Earth (volumetric mean)	$R_E$	$6371 \text{ km}$
Average radius of the Moon	$R_M$	$1740 \text{ km}$
Average radius of the Sun (visible disk)	$R_S$	$6.96 \times 10^5 \text{ km}$
Average solar flux at TOA	$E^0$	$1366 \text{ W m}^{-2}$
Mass of the earth	$M_E$	$5.988 \times 10^{24} \text{ kg}$

**TABLE A.3**  
Relevant Meteorological Constants

Density of air at standard pressure and temperature	$\rho$	$1.273 \times 10^{-3} \text{ g cm}^{-3}; 1.273 \text{ kg m}^{-3}$
Density of ice (0 °C)	$\rho_i$	$0.917 \times 10^3 \text{ kg m}^{-3}$
Density of liquid water (4 °C)	$\rho_l$	$1 \times 10^3 \text{ kg m}^{-3}$
Dry air gas constant	$R_{\text{air}}$	$287 \text{ J kg}^{-1} \text{ K}^{-1}$
Latent heat of fusion, ice	$L_{\text{ice}}$	$3.34 \times 10^5 \text{ J kg}^{-1}$
Latent heat of vaporization, water at 0 °C	$L$	$2.50 \times 10^6 \text{ J kg}^{-1}$
Molecular weight of dry air	$M$	$28.97 \text{ g mol}^{-1}$
Saturation vapour pressure (0 °C)	$e_0$	$6.1078 \text{ mb}$
Specific heat of air at constant pressure	$C_p$	$10.04 \times 10^2 \text{ J kg}^{-1} \text{ K}^{-1}$
Specific heat of air at constant volume	$C_v$	$7.17 \times 10^2 \text{ J kg}^{-1} \text{ K}^{-1}$
Standard pressure	$p_0$	$1013.25 \text{ mb } 1.013 \times 10^5 \text{ Pa}$
Standard temperature	$T_0$	$273.16 \text{ K}$

# B

## *Mathematical Definitions and Identities*

### **B.1 Mathematical Operators in Cartesian Coordinates**

del

$$\nabla f = \frac{\partial f}{\partial x} \mathbf{i}_x + \frac{\partial f}{\partial y} \mathbf{i}_y + \frac{\partial f}{\partial z} \mathbf{i}_z \quad (\text{B.1})$$

divergence

$$\text{div} f = \frac{\partial f}{\partial x} + \frac{\partial f}{\partial y} + \frac{\partial f}{\partial z} \quad (\text{B.2})$$

$$= \nabla \cdot f \quad (\text{B.3})$$

curl Laplacian

$$\nabla^2 f = \frac{\partial^2 f}{\partial x^2} + \frac{\partial^2 f}{\partial y^2} + \frac{\partial^2 f}{\partial z^2} \quad (\text{B.4})$$

### **B.2 Matrix Algebra**

For matrices **A** and **B**

$$(\mathbf{AB})^T = \mathbf{B}^T \mathbf{A}^T \quad (\text{B.5})$$

For a symmetric matrix **S**

$$\frac{\partial}{\partial \mathbf{x}} (\mathbf{x} - \mathbf{x}_a)^T \mathbf{S}^{-1} (\mathbf{x} - \mathbf{x}_a) = 2\mathbf{S} (\mathbf{x} - \mathbf{x}_a) \quad (\text{B.6})$$

$$\frac{\partial}{\partial \mathbf{x}} (\mathbf{y} - \mathbf{Kx})^T \mathbf{S}^{-1} (\mathbf{y} - \mathbf{Kx}) = -2\mathbf{K}^T \mathbf{S} (\mathbf{y} - \mathbf{Kx}) \quad (\text{B.7})$$

### B.3 Legendre Polynomials

The Legendre polynomials  $P_l(x)$ ,  $l = 0, 1, 2, \dots$ , are solutions of Legendre's differential equation

$$(1 - x^2) \frac{\partial^2 y}{\partial x^2} - 2x \frac{\partial y}{\partial x} + l(l + 1)y = 0 \quad (\text{B.8})$$

when  $x$  is a real number in the range  $[-1, 1]$ . For  $l \geq 0$  the polynomial can be written

$$P_l(x) = \sum_{r=0}^m \frac{(2l - 2r)!}{2^l r!(l - r)!(l - 2r)!} x^{l-2r} \quad (\text{B.9})$$

where the integer  $m$  is  $l/2$  or  $(l - 1)/2$ . The explicit forms of the first six Legendre polynomials are

$$\begin{aligned} P_0(x) &= 1, & P_1(x) &= x, \\ P_2(x) &= \frac{1}{2}(3x^2 - 1), & P_3(x) &= \frac{1}{2}(5x^3 - 3x), \\ P_4(x) &= \frac{1}{8}(35x^4 - 30x^2 + 3), & P_5(x) &= \frac{1}{8}(65x^5 - 70x^3 + 15x). \end{aligned} \quad (\text{B.10})$$

The polynomials can be evaluated at the specific arguments  $-1, 0$  and  $1$  to be

$$P_l(-1) = (-1)^l \quad (\text{B.11})$$

$$P_l(0) = \frac{\cos(l\pi/2)\Gamma[(l + 1)/2]}{\sqrt{\pi}\Gamma(l/2 + 1)} \quad (\text{B.12})$$

$$P_l(1) = 1 \quad (\text{B.13})$$

Recurrence relations can be used to evaluate higher order Legendre polynomials. Formulae include

$$P'_{l+1}(x) = (2l + 1)P_l(x) + P'_{l-1}(x) \quad (\text{B.14})$$

$$(l + 1)P_{l+1}(x) = (2l + 1)xP_l(x) - lP_{l-1}(x) \quad (\text{B.15})$$

$$P'_{l+1}(x) = xP'_l(x) + (l + 1)P_l(x) \quad (\text{B.16})$$

$$(x^2 - 1)P'_l(x) = lxP_l(x) - lP_{l-1}(x) \quad (\text{B.17})$$

$$P'_l(x) = \frac{l(l + 1)}{2l + 1} \frac{1}{1 - x^2} [P_{l+1}(x) - P_{l-1}(x)] \quad (\text{B.18})$$

where the prime denotes differentiation with respect to  $x$ . Finally the Legendre polynomials are interrelated through

$$P_{l-x} = (-1)^l P_l(x) \quad (\text{B.19})$$

$$P_{-l-1}(x) = P_l(x) \quad (\text{B.20})$$

## B.4 Associated Legendre Polynomials

The equation for the associated Legendre polynomials is [Abramowitz and Stegun, 1964]

$$(1-x^2)\frac{\partial^2 y}{\partial x^2} - 2x\frac{\partial y}{\partial x} + \left[l(l+1) - \frac{m^2}{1-x^2}\right]y = 0 \quad (\text{B.21})$$

has a solution (the solution needs a reference, also need to discuss arguments greater than 1)

$$P_l^m(x) = (1-x^2)^{m/2} \frac{d^m P_l(x)}{dx^m} \quad (\text{B.22})$$

where  $x$  is restricted to  $[-1, 1]$ . Some authors include a factor  $(-1)^m$  in the definition of  $P_l^m(x)$ . The associated Legendre polynomials have the following explicit forms

$$\begin{aligned} P_0^0(x) &= 1 \\ P_1^0(x) &= x & P_1^1(x) &= (1-x^2)^{1/2} \\ P_2^0(x) &= \frac{1}{2}(3x^2-1) & P_2^1(x) &= 3x(1-x^2)^{1/2} & P_2^2(x) &= 3(1-x^2) \\ P_3^0(x) &= \frac{1}{5}(5x^3-3x) & P_3^1(x) &= \frac{3}{2}(5x^3-1)(1-x^2)^{1/2} & P_3^2(x) &= 15(1-x^2) \end{aligned} \quad (\text{B.23})$$

also

$$P_l^0(x) = P_l(x) \quad \text{and} \quad P_l^{-m}(x) = (-1)^m \frac{(l-m)!}{(l+m)!} P_l^m(x). \quad (\text{B.24})$$

The associated Legendre polynomials obey the following recurrence relations

$$(2l+1)xP_l^m(x) = (l+m)P_{l-1}^m(x) + (l-m+1)P_{l+1}^m(x), \quad (\text{B.25})$$

$$(1-x^2)^{1/2} \frac{dP_l^m(x)}{dx} = \frac{1}{2}(l+m)(l-m+1)P_{l-1}^{m-1}(x) + \frac{1}{2}P_{l+1}^{m+1}(x), \quad (\text{B.26})$$

$$(x^2-1)P_l^m(x) = lxP_{l-1}^m(x) - (l+m)P_{l-1}^m(x). \quad (\text{B.27})$$

## B.5 Mie Angular Functions

The Mie angular functions  $\pi_n$  and  $\tau_n$  are defined in terms of the associated Legendre polynomials as

$$\pi_n(\cos \Theta) = \frac{1}{\sin \Theta} P_n^1(\cos \Theta) \quad (\text{B.28})$$

$$\tau_n(\cos \Theta) = \frac{d}{d\Theta} P_n^1(\cos \Theta). \quad (\text{B.29})$$

Evaluating the first terms gives

$$\begin{aligned} \pi_0(\cos \Theta) &= 0, & \pi_1(\cos \Theta) &= 1, & \pi_2(\cos \Theta) &= 3 \cos \Theta \\ \tau_0(\cos \Theta) &= 0, & \tau_1(\cos \Theta) &= \cos \Theta, & \tau_2(\cos \Theta) &= 3 \cos 2\Theta \end{aligned} \quad (\text{B.30})$$

$$\pi_n(\cos \Theta) = \frac{1}{\sin \Theta} \frac{(2n-1) \cos \Theta P_{n-1}(\cos \Theta) - (n-1)P_{n-2}(\cos \Theta)}{n} \quad (\text{B.31})$$

$$\frac{1}{\sin \Theta} \frac{(2n-1) \cos \Theta P_{n-1}(\cos \Theta)}{n} - \frac{(n-1)P_{n-2}(\cos \Theta)}{n} \quad (\text{B.32})$$

## B.6 Bessel Functions

The solutions to Bessel's Equation

$$x^2 \frac{d^2 y}{dx^2} + x \frac{dy}{dx} + (x^2 - n^2)y = 0 \quad (\text{B.33})$$

are called Bessel functions. The order of the Bessel function solution is determined by the constant  $n$ . When  $n$  is not an integer the two solutions of Equation B.33 are the Bessel function of the first kind of order  $n$ ,  $J_n(x)$ . Expressed as a series [see *Boas*, 2006, for the derivation] they are

$$J_n(x) = \sum_{m=0}^{\infty} \frac{(-1)^m}{\Gamma(m+1)\Gamma(m+1+n)} \left(\frac{x}{2}\right)^{2m+n} \quad (\text{B.34})$$

and

$$J_{-n}(x) = \sum_{m=0}^{\infty} \frac{(-1)^m}{\Gamma(m+1)\Gamma(m+1-n)} \left(\frac{x}{2}\right)^{2m-n} \quad (\text{B.35})$$

giving the general solution to Bessel's equation of

$$y(x) = c_1 J_n(x) + c_2 J_{-n}(x) \quad n \text{ not an integer} \quad (\text{B.36})$$

where  $c_1$  and  $c_2$  are constants.

To provide solutions at integer orders a combination of  $J_n(x)$  and  $J_{-n}(x)$  are used as the second solution of Bessel's equation. These are  $Y_n(x)$  called Bessel functions of the second kind of order  $n$  and defined

$$Y_n(x) = \frac{\cos \pi n J_n(x) - J_{-n}(x)}{\sin \pi n}. \quad (\text{B.37})$$

In some texts Bessel functions of the second kind are called the Weber functions or the Neumann functions (and so written  $N_n(x)$ ). The general solution of Bessel's equation applicable for all  $n$  is then

$$y(x) = c_3 J_n(x) + c_4 Y_n(x), \quad \text{all } n \quad (\text{B.38})$$

where  $c_3$  and  $c_4$  are constants.

Finally a complex linear combination of Bessel functions of the first and second kinds is used to form Bessel functions of the third kind (also known as Hankel functions)  $H_n^{(1)}$  and  $H_n^{(2)}$ . These are defined

$$H_n^{(1)}(x) = J_n(x) + iY_n(x), \quad (\text{B.39})$$

$$H_n^{(2)}(x) = J_n(x) - iY_n(x). \quad (\text{B.40})$$

*Boas* [2006] develops the following relations between the Bessel functions and their derivatives (which are also valid for the Neumann function),

$$\frac{d}{dx} [x^n J_n(x)] = x^n J_{n-1}(x), \quad (\text{B.41})$$

$$\frac{d}{dx} [x^{-n} J_n(x)] = -x^{-n} J_{n+1}(x), \quad (\text{B.42})$$

$$J_{n-1}(x) + J_{n+1}(x) = \frac{2n}{x} J_n(x), \quad (\text{B.43})$$

$$J_{n-1}(x) - J_{n+1}(x) = 2J'_n(x), \quad (\text{B.44})$$

$$J'_n(x) = -\frac{n}{x} J_n(x) + J_{n-1}(x) = \frac{n}{x} J_n(x) - J_{n+1}(x). \quad (\text{B.45})$$

## B.7 Spherical Bessel Functions

The spherical Bessel function are defined in terms the Bessel functions of half-odd interger order. They are [*Boas*, 2006]

$$j_n(x) = \sqrt{\frac{\pi}{2x}} J_{n+1/2}(x), \quad (\text{B.46})$$

$$y_n(x) = \sqrt{\frac{\pi}{2x}} Y_{n+1/2}(x), \quad (\text{B.47})$$

$$h_n^1(x) = j_n(x) + iy_n(x) = \sqrt{\frac{\pi}{2x}} H_{n+1/2}^{(1)}(x), \quad (\text{B.48})$$

$$h_n^2(x) = j_n(x) - iy_n(x) = \sqrt{\frac{\pi}{2x}} H_{n+1/2}^{(2)}(x). \quad (\text{B.49})$$

The power series expansions for the first and second order spherical Bessel functions are

$$j_n(x) = 2^n x^n \sum_{m=0}^{\infty} \frac{(-1)^m (n+m)!}{m!(2n+2m+1)!} x^{2m}, \quad (\text{B.50})$$

$$y_n(x) = -\frac{1}{2^n x^{n+1}} \sum_{m=0}^{\infty} \frac{\Gamma(2n-2m+1)}{m!\Gamma(n-m+1)} x^{2m}. \quad (\text{B.51})$$

Evaluating these expressions for the first few terms gives

$$\begin{aligned} j_0(x) &= 1 - \frac{x^2}{6} + \frac{x^4}{120}, & j_1(x) &= \frac{x}{3} - \frac{x^3}{30} + \frac{x^5}{840}, & j_2(x) &= \frac{x^2}{15} - \frac{x^4}{210} + \frac{x^6}{3780}, \\ y_0(x) &= -\frac{1}{x} + \frac{x}{2} - \frac{x^3}{24}, & y_1(x) &= -\frac{1}{x^2} - \frac{1}{2} + \frac{x^2}{4}, & y_2(x) &= -\frac{3}{x^3} - \frac{1}{2x} - \frac{x}{8}. \end{aligned} \quad (\text{B.52})$$

## B.8 Ricatti-Bessel Functions

The Ricatti-Bessel functions  $\psi_n(x)$ ,  $\chi_n(x)$  and  $\zeta_n(x)$  are the product of the relevant spherical Bessel function with its argument, i.e.

$$\psi_n(x) = x j_n(x) = \sqrt{\frac{\pi x}{2}} J_{n+1/2}(x), \quad (\text{B.53})$$

$$\chi_n(x) = x y_n(x) = \sqrt{\frac{\pi x}{2}} Y_{n+1/2}(x), \quad (\text{B.54})$$

$$\zeta_n(x) = \psi_n(x) + i \chi_n(x) = \sqrt{\frac{\pi x}{2}} H_{n+1/2}^{(1)}(x). \quad (\text{B.55})$$

The infinite series expressions are

$$\psi_n(x) = 2^n x^{n+1} \sum_{m=0}^{\infty} \frac{(-1)^m (n+m)!}{m! (2n+2m+1)!} x^{2m}, \quad (\text{B.56})$$

$$\chi_n(x) = -\frac{1}{2^n x^n} \sum_{m=0}^{\infty} \frac{\Gamma(2n-2m+1)}{m! \Gamma(n-m+1)} x^{2m}. \quad (\text{B.57})$$

Alternatively they can be written as a terminating series [Gumprecht and Sliepcevich, 1951]

$$\begin{aligned} \psi_n(x) &= \sin\left(x - \frac{n\pi}{2}\right) \sum_{m=0}^{\leq n/2} \frac{(-1)^m (n+2m)!}{(2m)! (n-2m)! (2x)^{2m}}, \\ &+ \cos\left(x - \frac{n\pi}{2}\right) \sum_{m=0}^{\leq (n-1)/2} \frac{(-1)^m (n+2m+1)!}{(2m+1)! (n-2m-1)! (2x)^{2m+1}}, \end{aligned} \quad (\text{B.58})$$

$$\begin{aligned} \chi_n(x) &= (-1)^n \cos\left(x + \frac{n\pi}{2}\right) \sum_{m=0}^{\leq n/2} \frac{(-1)^m (n+2m)!}{(2m)! (n-2m)! (2x)^{2m}}, \\ &- \sin\left(x - \frac{n\pi}{2}\right) \sum_{m=0}^{\leq (n-1)/2} \frac{(-1)^m (n+2m+1)!}{(2m+1)! (n-2m-1)! (2x)^{2m+1}}. \end{aligned} \quad (\text{B.59})$$

Using these expressions the Ricatti-Bessel functions for the first two orders are

$$\begin{aligned} \psi_0(x) &= \sin x & \chi_0(x) &= \cos x & \zeta_0(x) &= \sin x + i \cos x \\ \psi_1(x) &= \cos x & \chi_1(x) &= -\sin x & \zeta_1(x) &= \cos x - i \sin x \end{aligned} \quad (\text{B.60})$$



Higher order values can be calculated using the recurrence relation

$$\zeta_n(x) = \frac{2n-1}{x} \zeta_{n-1}(x) - \zeta_{n-2}(x) \quad (\text{B.61})$$



# C

## *Series Expansion of the Solution of the Spherical Wave Equation*

The three dimensional wave or Helmholtz equation

$$\nabla^2 \Pi + k^2 \Pi = 0 \quad (\text{C.1})$$

can be expressed in spherical coordinates as [Kerker, 1969]

$$\frac{1}{r} \frac{\partial^2 (r\Pi)}{\partial r^2} + \frac{1}{r^2 \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \Pi}{\partial \theta} \right) + \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 \Pi}{\partial \phi^2} + k^2 \Pi = 0 \quad (\text{C.2})$$

Using the method of separation of variables, a solution of the form

$$\Pi = R(r)\Theta(\theta)\Phi(\phi) \quad (\text{C.3})$$

is adopted. Substituting this into Equation C.2 and multiplying by  $r^2/R\Theta\Phi$  gives

$$\frac{r}{R} \frac{\partial^2 (rR)}{\partial r^2} + \frac{1}{\Theta \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \Theta}{\partial \theta} \right) + \frac{1}{\Phi \sin^2 \theta} \frac{\partial^2 \Phi}{\partial \phi^2} + r^2 k^2 = 0 \quad (\text{C.4})$$

which can be multiplied by  $\sin^2 \theta$  so that the third term is purely a function of  $\phi$ . Separating the equation gives

$$\left[ \frac{r}{R} \frac{\partial^2 (rR)}{\partial r^2} + \frac{1}{\Theta \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \Theta}{\partial \theta} \right) + r^2 k^2 \right] \sin^2 \theta = m^2, \quad (\text{C.5})$$

$$\frac{1}{\Phi} \frac{\partial^2 \Phi}{\partial \phi^2} = -m^2. \quad (\text{C.6})$$

The requirement of a single value solution implies that the solution for  $\Phi$  must have the same value for  $\phi$  as for  $\phi + 2m\pi$  where  $m$  is an integer. Applying this restriction means that the solution for Equation C.6 is

$$\Phi = a_m \cos(m\phi) + b_m \sin(m\phi). \quad (\text{C.7})$$

where  $a_m$  and  $b_m$  are constants.

Equation C.5 can be rewritten as

$$\frac{r}{R} \frac{\partial^2 (rR)}{\partial r^2} + \frac{1}{\Theta \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \Theta}{\partial \theta} \right) + r^2 k^2 - \frac{m^2}{\sin^2 \theta} = 0 \quad (\text{C.8})$$

so that by introducing a further constant  $p$ , it can be separated into

$$\frac{r}{R} \frac{\partial^2(rR)}{\partial r^2} + r^2 k^2 = p \quad (\text{C.9})$$

$$\frac{1}{\Theta \sin \theta} \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial \Theta}{\partial \theta} \right) - \frac{m^2}{\sin^2 \theta} = -p \quad (\text{C.10})$$

Substituting  $x = \cos \theta$  reveals that Equation C.10 is the equation for the associated Legendre functions [Abramowitz and Stegun, 1964]

$$(1 - x^2) \frac{\partial^2 \Theta}{\partial x^2} - 2x \frac{\partial \Theta}{\partial x} + \left[ l(l+1) - \frac{m^2}{1-x^2} \right] \Theta = 0 \quad (\text{C.11})$$

given that  $p = l(l+1)$  and  $l$  must be a integer in order that the solutions are finite at  $x = \cos \theta = \pm 1$ . The solution of C.10 is then

$$\Theta = P_l^m(\cos \theta). \quad (\text{C.12})$$

where  $P_l^m$  are the associated Legendre functions. They can be evaluated by means of recursion relationships defined in Appendix B.

The radial expression (Equation C.9) is now expressed as

$$\frac{\partial^2(rR)}{\partial r^2} + \left[ k^2 - \frac{l(l+1)}{r^2} \right] rR = 0 \quad (\text{C.13})$$

and solved by making the substitutions

$$kr = \rho \quad \text{and} \quad R(r) = \frac{1}{\sqrt{\rho}} Z(\rho) \quad (\text{C.14})$$

to obtain the Bessel equation of half integer order

$$\rho^2 \frac{\partial^2 Z}{\partial \rho^2} + \rho \frac{\partial Z}{\partial \rho} + \left[ \rho^2 - \left( l + \frac{1}{2} \right)^2 \right] Z = 0 \quad (\text{C.15})$$

The two solutions of this expression are the half integral order Bessel function  $J_{l+1/2}(\rho)$  and the half integral order Neumann function  $N_{l+1/2}(\rho)$ . The general solution of Bessel's equation [Boas, 2006] may be written as

$$Z(\rho) = A J_{l+1/2}(\rho) + B N_{l+1/2}(\rho) \quad (\text{C.16})$$

where  $A$  and  $B$  are arbitrary constants. The solution of Equation C.9 to be written as

$$R(r) = \frac{1}{\sqrt{kr}} [A J_{l+1/2}(kr) + B N_{l+1/2}(kr)] \quad (\text{C.17})$$

Choosing different constants  $c_l$  and  $d_l$  defined by

$$c_l = \frac{A}{k} \sqrt{\frac{2}{\pi}} \quad \text{and} \quad d_l = -\frac{B}{k} \sqrt{\frac{2}{\pi}} \quad (\text{C.18})$$

allows Equation C.17 to be expressed as

$$rR(r) = c_l \sqrt{\frac{\pi kr}{2}} J_{l+1/2}(kr) - d_l \sqrt{\frac{\pi kr}{2}} N_{l+1/2}(kr) \quad (\text{C.19})$$

$$= c_l \psi_l(kr) + d_l \chi_l(kr) \quad (\text{C.20})$$

where  $\psi_l(\rho)$  and  $\chi_l(\rho)$  are the Ricatti-Bessel functions which are defined in terms of the Bessel and Neumann functions [Kerker, 1969] by

$$\psi_l(\rho) = \sqrt{\frac{\pi \rho}{2}} J_{l+1/2}(\rho), \quad (\text{C.21})$$

$$\chi_l(\rho) = -\sqrt{\frac{\pi \rho}{2}} N_{l+1/2}(\rho). \quad (\text{C.22})$$

Hence

$$rR = c_l \psi_l(kr) + d_l \chi_l(kr) \quad (\text{C.23})$$

The Ricatti-Bessel functions can be evaluated by means of recursion relationships defined in Appendix B. The functions  $\chi_l(kr)$  become infinite at the origin so can not be used to represent a wave where  $r$  can equal 0.

The general solution of the wave equation in spherical coordinates is obtained from a linear superposition of all the particular solutions

$$\begin{aligned} r\Pi &= r \sum_{n=0}^{\infty} \sum_{m=-n}^n \pi_n^m \\ &= \sum_{n=0}^{\infty} \sum_{m=-n}^n [c_l \psi_l(kr) + d_l \chi_l(kr)] P_l^m(\cos \theta) [a_m \cos(m\phi) + b_m \sin(m\phi)] \end{aligned} \quad (\text{C.24})$$



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## *References*





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